JWST Master Class Workshop

Available proposal tools: MPT, ETC & JIST

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MSA Planning Tool (MPT)



A guide to JWST MOS terminology

- **MOS** Multi-Object Spectroscopy
 - the NIRSpec mode for obtaining spectra of 10-100s of specific targets simultaneously
- MSA Micro-Shutter Assembly
 - the array of very small shutters (0.2x0.46 arcsec) within NIRSpec that allows specific targets to be selected and contaminants to be blocked
- MPT MSA Planning Tool
 - the tool within APT for planning and optimising the MOS observations with the NIRSpec MSA



Why use MPT?

- NIRSpec MOS can obtain spectra for 10s-100s of objects simultaneously
- Targets are selected by opening a few of a grid of 250,000 micro-shutters
- MPT optimises this "MOS mask" given your target catalogue



- details on the MSA shutter array are in the "observing modes" session
- strategies are covered in the MOS hands-on session

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Example (flat-field test exposure) of 100 tightly-packed spectra (G395H grating)







MSA Planning Tool (MPT) in APT



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What is needed to run MPT?

- An internet connection
 - access to the most up-to-date MSA shutter operability
 - checking for guide stars during planning
- A complete and accurate astrometric catalogue • accurate source positions (<15mas relative accuracy)... may require NIRCam</p>
 - pre-imaging
- MPT produces:
 - plans with 1 or more pointings
 - MSA configurations
 - target sets
- One or multiple plans can be selected for a given observation





The catalogue

- First step to create a MOS observation... upload complete catalogue of sources to the MPT
- Catalogue should include all known sources in the field, to properly identify contaminants
- Upload source catalogue as an ASCII file
- It **must** contains J2000 **RA** and **Dec** expressed in degrees or hexadecimal
- It cannot have duplicate IDs or NULL entries
- Optional: Fluxes or magnitudes for the sources helps
 - Magnitudes are needed to properly define reference stars
- Optional: Target priorities are recommended
- The file can have a header, marked by "#"
- Remember: the relative astrometric accuracy of the catalogue must be between 5 and 50 mas

# ID	RA	DEC MA	F160W MAG	RR_F160V	FWHM	STELLARITY
514	53.17530756	-27.81989068	22.46	0.039	2.9	0.91
2639	53.17797177	-27.80327718	99	29.183	3.48	0.81
7894	53.16615975	-27.76428237	29.284	0.202	6.95	0.81
3352	53.15236091	-27.79958882	29.388	0.275	6.6	0.74
10101	53.18123196	-27.78147671	28.977	0.323	5.29	0.74
4166	53.16562375	-27.79587255	28.733	0.136	9.86	0.73
6093	53.15093966	-27.79386221	30.681	0.576	3.26	0.73
7740	53.16000591	-27.76349436	29.364	0.218	8.99	0.73
9976	53.14403028	-27.78036021	29.263	0.206	7.82	0.73
3740	53.16054039	-27.79774121	29.213	0.228	6.21	0.71
10586	53.1390074	-27.78953776	29.582	0.268	9.7	0.71
615	53.15839939	-27.81899646	24.087	0.06	2.84	0.7
8694	53.17044959	-27.77458204	29.424	0.38	5.34	0.7
2032	53.14225186	-27.80676448	29.2	0.636	9.56	0.69
6456	53.16153898	-27.78625609	29.181	0.189	10.11	0.69
7919	53.16679656	-27.76437467	29.013	0.289	7.18	0.68
4567	53.12870239	-27.78690427	99	28.452	4.67	0.67
7988	53.16384488	-27.76478608	28.148	0.277	4.5	0.66
5914	53.16015634	-27.79318966	29.424	0.25	8.94	0.64
20309	53.16691677	-27.81874945	27.564	0.118	3.41	0.63
8737	53.15973205	-27.76936156	29.291	0.218	3.08	0.62
7566	53.18247561	-27.78085107	29.067	0.21	6.28	0.58
9681	53.18741752	-27.77796957	29.419	0.289	3.24	0.55



Load the catalogue as an MSA Source Catalogue Target



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Catalogue importer

• Column for Flux can be used for filtering sources when creating candidate sets • Weight can be used to prioritise targets

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ID	RA	DEC	Ignore Colui	Ignore Column	FWHM	Stellarity	Ignore Column	Ignore Col	Ignore Col	Ignore Coli	Ignore Coli	Ignore Coli	Ignore Colum	Ignore Col	Weight	Magnitude	Magnitude	Magnitude
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Cancel	X Import																	

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MSA Source Catalogue Target

Number 1



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The MSA Planning Tool

MSA Planning Tool in the APT toolbar

The **Planner** tab is where you design MOS plans

- Select candidate lists from your MSA Catalog
- Aperture PA (position angle) is either a placeholder, or assigned to you by STScl (after acceptance)
- Slitlet configuration and Centering **Constraint** should be chosen



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The MSA Planning Tool

The Planner is where you decide how to **dither**

- **Nodding** moves the sources within the slitlet – no shutter reconfiguration
- **Fixed Dither** moves the sources by a finite number of shutters specified by the user along the dispersion and/or the crossdispersion direction.

Grating/Filter combination(s) must be selected to prevent spectral overlap in the chosen configurations



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Planner Timeline	🧭 🔶 iew in Aladin 🛛 BOT	Target Confirmation	PDF Preview	Submission I	Frrors and Warning	5		Run All Tools	Stop
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sts									
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	G395M/F290LP								
	PRISM/CLEAR								
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Center	RA: 03 32 39.6524	Dec: -	27 47 26.91						
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Examine a

The **Plans** tab is where you examine and visualise MOS plans

- MSA shutter view
- Collapsed shutter view
- Send to Aladin (onsky visualisation)

nd vis	▲ Astronomer's Proposal Tools Version 2020.1 mpt-demo (Thu Oct 24 2019) - JWST Draft Proposal (RAFEL-2015.aptx)	
1	\gg \mathbb{E} \mathbb{E} \mathbb{E} \mathbb{E}	
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Examine and visualise a plan

Targets on the MSA

	Another Plan c1e1
Image: constrained of the second of the se	Ready
	Save as svg Export to CSV PIPPO (9969 sources)
	Add s

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Shutters projected onto the sky in Aladin



hutters plane to Aladin



Create an observation

From the **Plans** tab

- Highlight a plan or plans
- Select the pointings
- Create Observation

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SS MASTER CLAS



Proposing for MOS

- New version of the MPT available for the Cycle 1 call • Remember: need a catalogue with accurate relative astrometry
- - <15mas relative accuracy</p>
- MOS will have a multi-step planning process
 - Final plans not possible until the program is accepted and scheduled (giving the actual Aperture PA)
 - A particular APA can be requested in a proposal, but must be strongly justified
- Instrument overheads are not negligible use APT/MPT for the estimate



Additional helpful hints

- **Dither!!!** ... to improve background subtraction, wavelength coverage, etc
- Order your input catalogue by weight... during optimisation, MPT adds sources in the catalogue order
- Use a **Filler Set** to maximise efficiency/multiplexing
- Include **Primary** candidates in the **Filler** list to obtain additional observations of important sources
- Weight only matters for **Primary** sources (not for **Fillers**)
- If observability window is large, test optimisation for several APAs... if multiplexing differs significantly, consider including an Orient Special Requirement, with a minimum range of 30 deg
- If using a high-res grating (GxxxH), attempt to place most sources on the leftmost quadrants (MPT orientation) to avoid detector cutoffs
- Use Aladin to show the position of NIRCam parallels during NIRSpec exposures
- Input catalogue should be as complete as possible to check for contaminants in commanded open (or known Failed Open) shutters
- The MSA Configuration Editor can be used to amend MPT-optimised configurations. It can be found at the **observation level** in APT



Exposure Time Calculator (ETC)



Where is ETC?

jwst.etc.stsci.edu

Welcome to the JWST Exposure Time Calculator

Login

Quick Start

Create User

News

Welcome to version 1.5 of the JWST ETC!

This release features new instrument modes, accuracy improvements, usability enhancements, and more: see the Release Notes for details, and be sure to review the Known Issues for this release.

When you log in to the 1.5 ETC, your old workbooks will be marked "Out of Date":

- When you load them, they will open in Read-Only mode: this ensures that your previous results are not overwritten and remain available to you for reference.
- If you copy an out of date workbook, and load the copy, all its calculations will be automatically updated for you with the current version of the software.
- For more information, see ETC Releases and Out-of-Date Workbooks.
- Pontoppidan et al., Proc. SPIE. 9910, Observatory Operations: Strategies, Processes, and Systems VI, 991016. July 15, 2016 (https://arxiv.org/abs/1707.02202)

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Work Anonymously

The "Pandeia" project

3D ETC engine Python library

Web application (jwst.etc.stsci.edu)

User interface, collaborative function

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Reference database

Throughputs, detector noise, PSFs, ...

JWST background model

Integrated into web application or as a standalone tool

The engine algorithm





Telescope

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Instrument

Detector

JWST ETC concept

Modern design

- Signal (source + background) modelled in 3D
 - conserves flux
 - resolve lines
 - oversample JWST resolution
- Pixel based:
 - Models the detector (e.g. correlated noise)
 - Final S/N calculation includes data analysis and post-processing steps ("strategy")
 - Allows the modelling of complex scenes (e.g. estimate contamination from bright sources)

Supports all JWST modes









PSF library

- Uses WebbPSF to calculate PSF including realistic wavefront errors
- Almost 5000 individual monochromatic PSFs
- Subsampled by integer factor of pixel size



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cluding realistic wavefront errors natic PSFs

NIRCam SW PSFs (0.6-5.2 micron)

Readout terminology

- subarray: the window on the detector being read out
- frame: one complete read of the detector or subarray
- **group**: a set of consecutive frames averaged onboard (some frames may be skipped)
- integration: one non-destructive ramp
- **exposure**: a set of consecutive integrations at the same pointing
- dither: an exposure at a new pointing



Time



Time terminology

- exposure time: Time the detector is operating during a single exposure
 - includes resets
 - excludes initial synchronisation time
 - Includes all integrations, but no repetitions per pointing, per tile, or per observation
- measurement time: For any individual pixel, the interval between first and last measurements during a single integration, multiplied by the number of integrations per exposure
 - used to determine count rate
- saturation time: For any individual pixel, the interval from reset to the final read of an integration, multiplied by the number of integrations per exposure
 - depends on exposure parameters, not target brightness or instrument throughput
- exposure duty cycle: measurement time divided by exposure time



How is saturation treated in ETC?



- Two types of saturation
 - Partial saturation: the integration saturates before it completes, but more than the minimum number of groups are unsaturated - DATA RECOVERED
 - Full saturation: the integration saturates before the minimum number of groups achieved - NO DATUM (for that pixel)
 - Usually, the minimum ngroups = 2





Tips for optimising detector set-up

- More frames per integration decreases read noise
- Longer groups decreases data volume
- Shorter groups decreases chance of cosmic ray hit per group
- More groups per integration make the cosmic ray correction better
- Longer integrations may make ramp fitting more certain
- More dithered exposures decreases flat-field errors and the impact of bad pixels

Note: in ETC, all exposures are assumed to be dithers, except for time-series observations
 Readout patterns that skip lots of frames have higher read noise,

 Readout patterns that skip lots of fr but slightly better duty cycle



JWST ETC features

Workbooks

- Organise several ETC calculations into a workbook
- A workbook can contain multiple sources, scenes and calculations
- Workbooks remain in your MyST account, and are shareable

Reusability

- Calculations in a workbook start with reasonable defaults
- Simply copy a calculation, modify the inputs as desired and recalculate

Batch expansion

- Efficiently run a batch of many calculations
- Calculation is duplicated N times varying only the selected parameter
- Expansion over e.g. Ngroups offers a way to show the behaviour of SNR as a function of "exposure" time

Auto-update

Changes made to the input, flow through to associated forms and calculations

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Scenes and sources

- Scene: small postage-stamp of the sky (a few arcsec on a side)
 - A scene can have no source (just background) or multiple sources
 - Scenes can be reused in multiple calculations

• Source: specify SED, normalisation, extinction, emission lines and shape

- Each sources can be reused in multiple scenes
- Location: x,y offset (with orientation) within the postage stamp scene
- Shape: point source or extended (flat, Gaussian, Sersic, power law...)
- Spectrum:
 - Continuum: flat, black body, power law, templates
 - Lines: centre, width, flux
 - Normalise to magnitude/flux in JWST or HST bandpass, or at a λ
 - Upload a spectrum (ASCII or FITS format)

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Scene Sketch





Issues to think about when using ETC

• ETC is a highly versatile tool

• There are many parameters affecting ETC sensitivities and program preparation

- Different read-out patterns change the read noise
- Be aware of saturation
- Which background subtraction scheme?
- Is the background correct for your target?
- What extraction aperture is optimal? (Point or extended source?)
- Does your extraction aperture contain contaminating flux from other sources in the scene?

Remember: ETC approximates our current best knowledge and understanding of performance. There are remaining uncertainties associated with system throughputs, detector noise properties, etc., which will remain unknown until JWST is in flight.

BEWARE: The ETC is not intended to be a complete observation simulator

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sitivities and program preparation

r extended source?) nating flux from other sources in the scen

Useful links

- JWST ETC (v1.5)
 - <u>https://jwst.etc.stsci.edu/</u>
- **Documentation**:
 - ETC introduction
 - Links to all pages, video tutorials and more
 - https://jwst-docs.stsci.edu/jwst-exposure-time-calculator-overview
- ETC release notes, known issues and FAQ (@ JWST Help Desk)
 - JWST Help Desk ETC FAQ
 - JWST Help Desk ETC known issues
- JWST Community Lecture Webcasts
 - <u>"The JWST Exposure Time Calculator"</u> Klaus Pontoppidan (Nov 2019)
 - <u>"JWST ETC Demo" (interface demonstration)</u> Swara Ravindranath (Feb 2017)

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- Quick-look observation feasibility
- Runs directly in a browser window
 - Online tool: <u>jist.stsci.edu</u> JDox: JIST
- For all basic observing modes, explore S/N values by adjusting source flux or telescope exposure parameters

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• Example view of the interface



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- Beware of the simplifying assumptions...
 - All calculations based on a point source:
 - Flat SED
 - Background spectrum set to 120% of minimum zodiacal background at a reference point (see usage notes for more detail)
 - Single integrations are considered. For multiple integrations (or dithers), SNR can be scaled by SQRT(NINTS)
 - Saturation simply handled by setting SNR=0 (JIST is not recommended) for saturated data)
 - For WFSS modes, only R Grism values presented (C Grism assumed) identical)





Accessing the tool - reminder



Accessing the tools

- Astronomer's Proposal Tool (APT)
- MSA Planning Tool (MPT)
 - \rightarrow
- Exposure Time Calculator (ETC) \rightarrow <u>jwst.etc.stsci.edu</u>
- JWST Interactive Sensitivity Tool (JIST) \rightarrow jist.stsci.edu

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apt.stsci.edu