JWST Master Class Workshop

NIRSpec Slit hands-on

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NIRSpec and MIRI LRS Slit hands-on

Infrared spectroscopy of Y dwarfs NIRSpec

MIRI LRS

- In this hands-on session you will be asked to use: The Exposure Time Calculator (ETC) for the NIRSpec
 - The Astronomer's Proposal Tool (APT) for NIRSpec

The MIRI LRS science case may be an extension to this exercise.



Thinking about strategies



NIRSpec Subarrays and Readout patterns

Subarray:

- FULL
- Tailored subarrays per individual slit: SUBS200A1, SUBS200A2, SUBS200B1, SUBS400A1
- S1600 subarrays: SUB2048, SUB1024A, SUB1024B, SUB512
- ALLSLITS (S200A1 and S200A2 combined to bridge the detectors gap, but not only. It can also be used to estimate background)

Fixed Slits (always open)	
Detector NRS1	

- FULL: only traditional readout mode
- **Readout patterns:** NRS, NRSRAPID, NRSIRS2, NRSIRS2RAPID
- Note: maximum exposure duration is 10,000 seconds

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Subarray name	Apertures available	Height (pixels)	Width (pixels)	Frame time (s)
FULL [†]	All FSs	2048	2048	10.73677 14.58889 (IRS2)
ALLSLITS	All FSs, option: <i>S200A1</i> and <i>S200A2</i> **	256	2048	5.49400
SUBS200A1	S200A1	64	2048	1.55800
SUBS200A2	S200A2	64	2048	1.55800
SUBS200B1	S200B1	64	2048	1.55800
SUBS400A1	S400A1	64	2048	1.55800
SUB2048	S1600A1	32	2048	0.90200
<i>SUB1024A</i> ^{‡*}	S1600A1	32	1024	0.45144
SUB1024B	S1600A1	32	1024	0.45144

32

512

Table 1. NIRSpec subarrays for the FS observing mode

S1600A1

SUB512



NIRSpec FS detector recommendations

Decision flow to specify detector parameters for a FS observation:





Background removal strategies

Pixel-to-pixel subtraction

- Compact sources: using nodding
- - exposures

Master background subtraction

- Standard for extended sources
- created from:

 - list of background elements (like off-source pixels)
- for NIRSpec: is set by pipeline processing if the grating wheel has moved
- In ETC, FS nod/dithers are accounted for at detector level, setting the number of exposures.
- interruptible

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Extended sources: associated dedicated "blank sky" observation at off-scene position - for NIRSpec: performed only if grating wheel has not moved between the target and off-scene

- associated exposure, creating an independently flux-calibrated 1D background spectrum

APT: Dedicated observations for background must be linked to science observations in order to create an association list. This is defined at Target Level in APT. Necessary for pipeline processing flow. APT: If the background signal is time variable throughout the year, dedicated "blank-sky" observations should be scheduled consecutively to ensure pixel-to-pixel subtraction. Special requirements: Non-

NIRSpec Nods and dithers

- Nods: offsets that produce data to be subtracted in pipeline processing, in order to cancel in-field background flux. Nod options are typically best used for targets that are not significantly spatially extended. (2, 3, or 5 points), depending on the number of exposures needed/possible.
- **Dithers:** offsets of the target position over multiple exposures, to even out or mitigate detector effects, help remove cosmic rays, improve spatial sampling, and increase signal-to-noise and flux accuracy.

- APT
 - **Primary dithers:** nod along the slit used to subtract background flux. Recommended 3, 5.
 - Sub-pixel dithers: to improve spatial and/or spectral pixel sampling, only in addition to Primary Dither pattern for point sources. Not recommended in spectral dimension.
 - Across gap SAM. Offset executed by selecting the option "S200A1 and S200A2" only in the high resolution gratings

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NONE SPECTRAL SPATIAL BOTH Subpixel positions = A A 8 8 000 **Primary nod** position Options = None, 2, 3, 5 B C 8 8 8 000 D B S 8 # of Exposures: 1 2 3 5 15 10 12 20





NIRSpec Nods and dithers

Decision tree for choosing the right nod/dither pattern for FS observations based on the source compactess.





NIRSpec FS Target Acquisition

FS TA methods:

- depends on centroiding accuracy of the target (ephemeris).
- depending on the catalogue relative accuracy.

Strategies and parameters:

- WATA:
 - Subarray configurations: SUB32, SUB2038, FULL (increasing frame time)
- MSATA:
 - Subarray configuration: FULL
- Filters: F110W, F140X, CLEAR
- Readout pattern: NRSRAPID, NRSRAPIDD1, NRSRAPIDD2, NRSRAPIDD6
- Groups/Integrations are fixed
- TA readout mode switch with respect to science parameters costs extra in time.

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WATA (recommended) using science or offset target centred in S1600A1. 11-18 minutes. Expected accuracy: 20 mas, and

MSATA requires defining 5-8 reference stars (may require pre-imaging). It is specified at the visit level of the observation, not directly at the observation template. 24-30 minutes. Expected accuracy: 20 – 25 mas (optimal), <50 mas (relaxed),

NONE is not recommended for FS. The resulting pointing accuracy will be that delivered by the GS acquisition at the start of the observation. For reference, the absolute pointing performance of JWST for NIRSpec is expected to be 100 mas









Science Overview

- Goal
 - To obtain spectroscopic observations of a Y dwarf across the entire JWST NIRSpec and MIRI LRS wavelength ranges to understand whether these atmospheres are shaped by chemical disequilibrium driven by vertical transport or the formation of water clouds, and constrain the object's gravity, hence mass.





Observations Methodology

Methodology

- Compare high-quality low and medium resolution Infrared spectra from 0.6 to 13 clouds.
- **Planned observations**
 - NIRSpec fixed slit spectroscopy
 - MIRI LRS slit spectroscopy
- Source Type
 - Point source
- **Observation strategy**
 - and MIRI LRS R~100
 - 3-point nod NIRSpec / 2-point nod MIRI
 - TA on science source

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microns, to models of cool atmospheres at different temperatures, gravity, degrees of turbulence, chemical equilibrium or disequilibrium driven by vertical transport, and

PRISM and G395M dispersers for NIRSpec to obtain R~100 and R~1000, respectively

Getting started ETC NIRSpec WB ID: 30896 MIRI LRS WB ID: 30989



Scenes and Sources

- NIRSpec:
 - 1 source with user supplied model spectrum morley_spec_ETC_noscale.txt renormalized to measured Vega magnitudes HST/ WFC3 F140X.
- MIRI
 - Workbook: 30989





Calculations







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ectra Caveats and Limitations

0					
	λ-	Scn -	(s) -	SNR -	A
	1.27	1	42.95	21.60	0
	1.00	1	1318.19	25.24	0
	4.70	1	5618.27	103.62	0
		-			-



NIRSpec Calculation for Target Acquisition

• WATA in ETC is available.





Scene 🖈 Backgrounds Inst	Trument Setup Detector Setup St	rategy	
Subarray	R	eadout pattern	
FULL	•	NRSRAPID	•
Groups 😯	Integrations	Exposures	
3	• 1	• 1	•
Total exposure time: 00:00:43 (42.	95 s)		
Total integrations: 1			





Cheat APT screenshots...



APT Fixed-slit spectroscopy templates

When creating a new observation, the user defines the AP

Number	4 Status:	Duplication
Label		
🗙 Instrument	None Selected	
Template	None Selected	\$

✓ NIRSpec Fixed Slit Spectroscopy NIRSpec IFU Spectroscopy NIRSpec MultiObject Spectroscopy NIRSpec Bright Object Time Series

- APT fixed-slit spectroscopy templates are divided into four sections:
 - Generic information
 - Observation information
 - Target acquisition parameters
 - Science parameters



Т	template
- C	

MIRI Imaging
✓ MIRI Low Resolution Spectroscopy
MIRI Medium Resolution Spectroscopy MIRI Coronagraphic Imaging



APT Fixed slit spectroscopy template

• NIRSpec

					👁 nirs
	Number	1 Status: U	NKNOWN		Duplication
	Label	nirspec-fs			
	Instrument	NIRSPEC			
	Template	NIRSpec Fixed Slit Sp	ectroscopy		
	Target	1 WISE-J035000.32-5	565830.2		
(Splitting Distance	Number of Visits		
\	/isit Splitting:	45.0 Arcsec	1	·	
		Science	Total Charged		
D	uration (secs)	6921	13178	r	
	Data Volume	2668 MB			

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spec-fs (Obs 1) of JWST Draft Proposal

Generic information



Observation information



APT Fixed slit spectroscopy template

• NIRSpec

				NIRS	pec Fixed S	lit Spectroscop	y Mosaic P	roperties	Special	Requirements Co	mments		
TA Method	WAT	Α	0										
	. D		.										
Target Acquisition	n Para	ame	ters										
			Acq Target		Acq Su	barray Acq	Filter						
Target ACQ	Sam	e Tar	get as Observatio	on ᅌ	⇒ FULI	L ᅌ F	140X ᅌ						
			Acq Readout Patt	ern	Acq Gr	oups/Int Acq	Integrations/Exp	Acq Total I	ntegration	s Acq Total Exposure Ti	me Acq ETC Wkbk.Ca	alc ID ETC	
Acq Exposure Time	NRS	RAPI	D		3	1		1		42.947	30896	ď	
Science Parameter	rs												
Slit	S20	0A1	• •										
Subarray	SUB	s200/	A1 🗘										
	Prin	nary I	Dither Positions	Sub-Pixe	el Pattern								
Dither Parameters	3			NONE	\$								
		#	Crating /Filter	Peadou	t Pattorn	Croups (Int	Integration	c/Evp A	utocal	Total Dithers	Total Integrations	Total Exposure Time	FTC Whele Cale ID
		1	PRISM/CLEAR	NRSRA	PID	140	2	S/EXP AL	ONE	3	6	1318,191	30896
	0	2	G395M/F290LP	NRSRA	PID	600	2	N	ONE	3	6	5618.271	30896
Gratings/Filters	C												
								Dualizata	Innet	Abaua			
							Add	Dupricate	Insert	Above Kemove			



Mosaic Properties Special Requirements Comments	
ations/Exp Acq Total Integrations Acq Total Exposure Time Acq	ETC Wkbk.Calc ID ETC
42.947 308	96 🕜

 ETC 2 2

APT Fixed slit spectroscopy template

• MIRI

	MIRLLow Resolution Spectroscopy	Mosaic Properties Spec	al Requirements Comments	
	Mixi Low Resolution Spectroscopy	Mosare Properties Spec	ar requirements comments	
Target Acquisition Parameters				
Acq Target	Acq Filter			
Target ACQ Same Target as Observation	n ᅌ 🔿 F560W ᅌ			
Acq Readout Pattern Acq G	Groups/Int Acq Integrations/Exp	Acq Total Integrations Acq Total E	xposure Time Acq ETC Wkbk.Calc ID	ETC
Acq Exposure Time FASTGRPAVG ᅌ 8	2	1 88.801	30989	
LRS Verification Image				
Obtain Verification Image? 🛛 Yes 💿 No				
LRS Parameters				
Subarray FULL ᅌ				
Dither Type				
Dither ALONG SLIT NOD ᅌ				
Readout Pattern Groups/Int In	ntegrations/Exp Exposures/Dith Total D	Dithers Total Integrations Total Exp	osure Time ETC Wkbk.Calc ID ETC	
Exposure Time FAST ᅌ 120 1	0 2 2	40 13320.1	30989)









Slit Help and JDOX

- **Detector Strategies**
 - NIRSpec detector recommended strategies
- NODS and Dithers
 - NIRSpec Background Recommended Strategies
 - MIRI LRS Recommended Strategies Background observations
 - NIRSpec Dithering Recommended Strategies FS
 - MIRI LRS Dithering Recommended Strategies
- Target Acquisition
 - NIRSpec Target Acquisition Recommended Strategies
 - MIRI Target Acquisition Generic Recommended Strategies
 - MIRI LRS Target Acquisition Recommended Strategies

